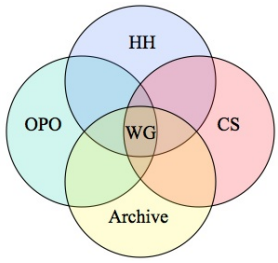


Broadening the Hubble Audience Base

BASIC TENET: "It should be possible to combine the strong public response to the **Hubble Heritage** images, the extensive educational expertise/infrastructure/audience of **OPO**, and the potential for **citizen science** embodied by the **new generation of archival products/tools** (e.g., galaxy zoo), to engender a **new level of interaction between Hubble and the community.**"

Three pronged approach:

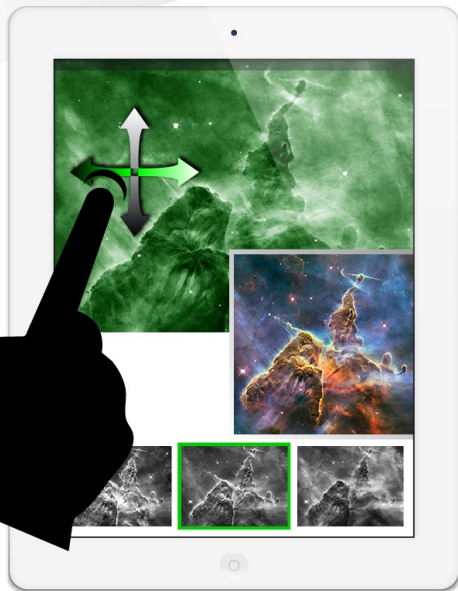
1. Astro Imaging – (Stratis)
2. Hubble Heritage / New Observations– (Zolt)
3. Working Group / Citizen Science – (Brad)



Astro Imaging

THE ARCHIVE FOR EVERYONE

PUBLIC IMPACT



Novice

DATA FOR
EDUCATION &
HOBBYISTS



Baby HLA






How Tos




Curriculum Support
Tools

Interested &
Directed

HUBBLE IMAGE PROCESSORS
RAIDERS OF THE
HUBBLE ARCHIVE

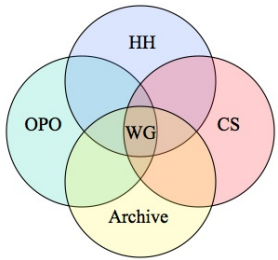




Hubble Legacy
Archive (HLA)



Multimission Archive
at Space Telescope
(MAST)

Expert



Hubble Heritage / New Observations

1. Aesthetic Hubble Images

Similar to past Heritage observations
Minimal orbits per target

2. Hubble Legacy Image

Substantial number of orbits for signature Hubble image.

Science involvement - include in yearly Call for Proposals

Likely to be large number of science papers (see backup).

3. “Other” Images – Targets of Opportunity (e.g., Mars), Anniversary, etc.

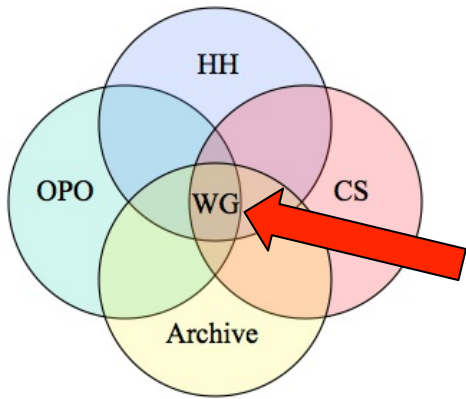
Timeline:

Now: Define targets/activities for ~15 orbits based on HH list.

In ~3 months: Devote ~15 orbits this year (probably Hubble Legacy Image) based primarily on outreach

Next Fall: Meet as group in the fall to select next years targets.

Notes: Always include High Level Science Products (HLSP) in archives at time of release.

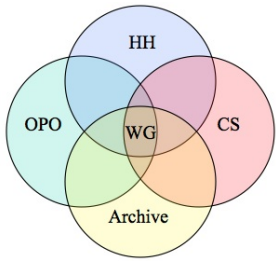


Working Group / Citizen Engagement Initiative

- A multi-division working group (including 3 - 4 members from Archives/ Citizen science, HH, OPO, and other interested staff) meet monthly for next 6 months.

Primary Initiatives:

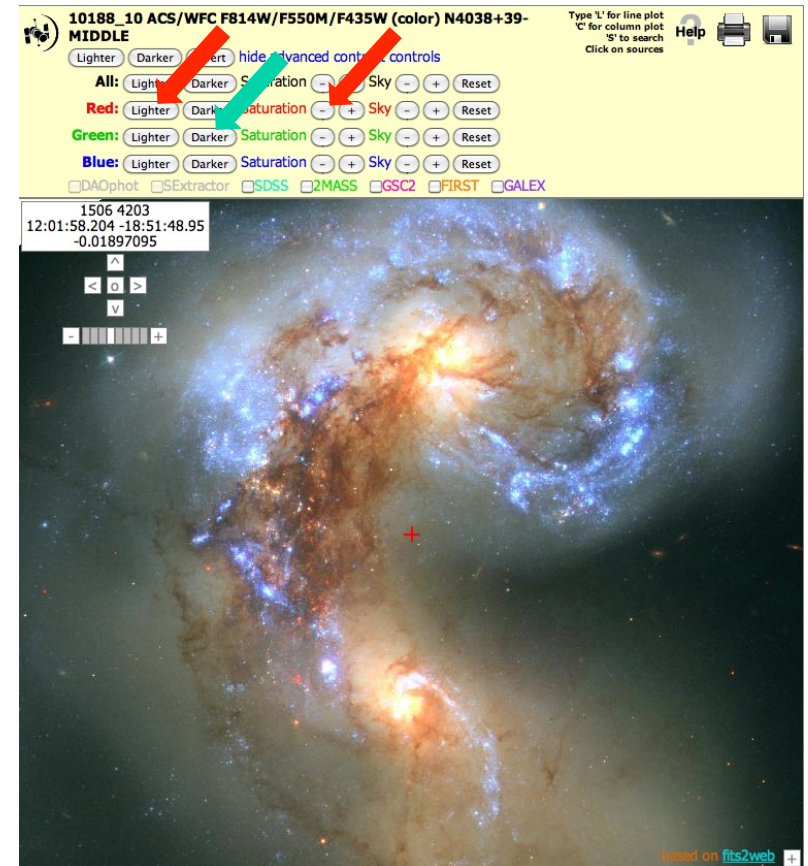
1. Citizen Science activities (e.g., Zooniverse, planet finding, age dating star clusters, ... – [see backup](#))
2. Archive lab exercises ([see backup](#)) – Often **requested by educators at HLA demos at AAS meetings**
3. Jamboree (solicit citizen science projects, EPO grants attached to successful proposers, ...)



Backup - Broadening the Hubble Audience

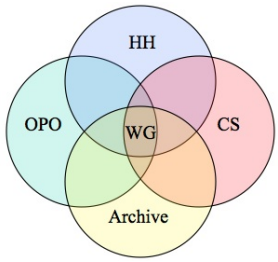
(longer version of introductory viewgraph)

- BASIC TENET: Citizen Science / Amateur Astronomy / Hands-on-Education represents a tremendous, largely untapped resource to engage the community and to engender their continued interest in Hubble (and JWST).
- PARAPHRASE OF STATEMENT FROM DIRECTOR: "It should be possible to combine the strong public response to the Hubble Heritage images, the extensive educational expertise/infrastructure/audience of OPO, and the potential for citizen science embodied by the new generation of archival products/tools (e.g., galaxy zoo), too engender a new level of interaction between Hubble and the community."

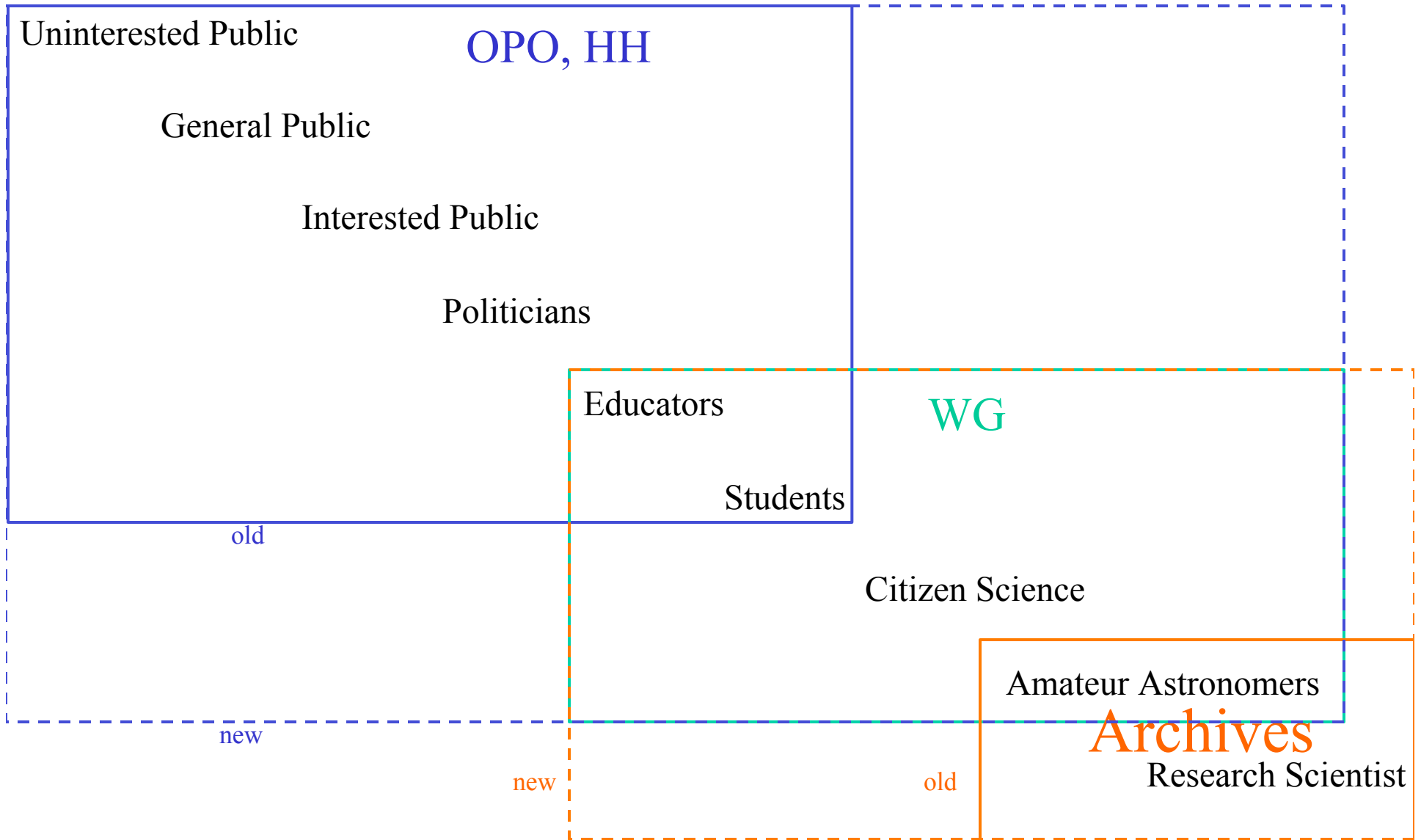


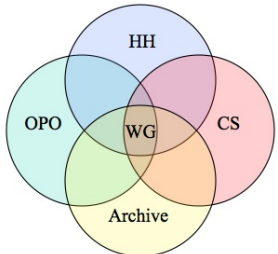
From HLA tutorial on how to make your own color images and submit to Hubble Hidden Treasure contest >>> [Demo](#)

http://hlatest.stsci.edu/iotm/number9_advanced_contrast_feb_22_2012.html



Backup: Who is the audience ?

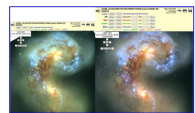




Backup: Past/Present/Future ASTRO IMAGING Activities



Image of the Month for March 2012
Advanced Contrast Controls / Hubble Hidden Treasures Contest



Highlighted HLA Features:

- Making your own color images using the Advanced Contrast Controls using NGC 4038+4039 (the "Antennae Galaxies") as an example.

Introduction:
Color images are automatically produced in the HLA whenever images are available in two or more different filters. Default values are used to set the brightness, saturation, and background key values for the red, green, and blue images which are used to make the color image. However, you can adjust these values yourself using the "tabernical control controls". You might also be interested in entering the "Hubble Hidden Treasures" contest described at "

This display can be reproduced in the HLA by using the following steps or clicking on the image above.

- Type [HLA.html](#) into your browser. Click on the "Enter Site here" button.
- Type "Antennae" in the search box and press the "return/enter" key. You will find that 379 different



ESO's Hidden Treasures (2010?)



Hubble Hidden Treasures 2012 contest (ESA lead, collaboration with OPO, HH, archives)



Focal point: Astro imaging website/facebook/flicker



In development: "Hubble Astro Imagers Handbook", starting with HLA tutorial and several documents Zolt has authored in past (Zolt, Brad, summer student?)



Hubble Heritage challenge



Planetary nebula challenge – "How do astronomers know what elements are present in astronomical objects ? > education activity

(i.e., **blue** = Oxygen, **green** = Hydrogen, **red** = Sulphur")

Hubble Hidden Treasures 2012

minisite text

Title/logo: Help us find Hubble's hidden treasures

What is it? [front page]

Hubble's Hidden Treasures is a competition which invites you to dig into Hubble's archive to find and process beautiful astronomical images.

There are two categories you can enter:

- Finding and processing Hubble observations using a set of simple online tools that anyone can learn;
- Finding and processing Hubble observations using professional tools.

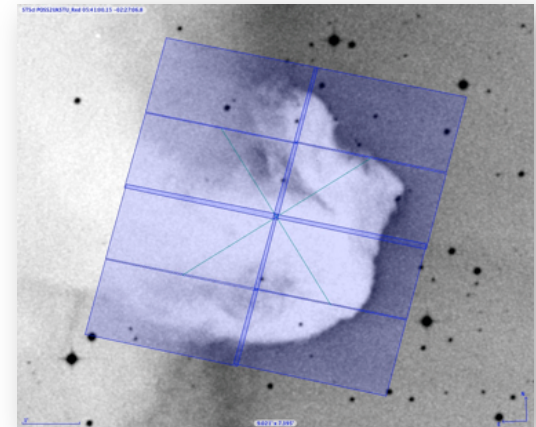
In both cases, you'll be playing with real data from the world's most famous astronomical observatory. You'll be helping to uncover hidden treasures from Hubble's archive for the whole world to enjoy. And you'll be in with a chance of winning one of our great prizes.

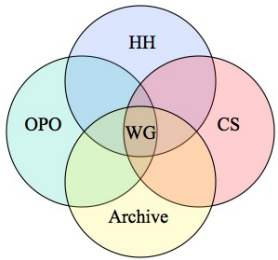
How does it work?

Dig through Hubble's science archive and try to find an amazing image that hasn't yet been published. Process the image, either using the simple online tools provided, or using professional astronomical image processing software. Then submit it to our Flickr group for a chance to win some great prizes [link](#). There is no limit to the number of entries per person.

Submit them to our Flickr group for a chance to win some great prizes [link](#).

This page gives a tutorial [link to tutorial](#) for how to search the Hubble archive. This page [link to 'What is image processing and how do I do it'](#) explains what image processing is and how you can do it, both with the simple online tools provided, or, if you want a challenge, using the same software as the professionals do.





Backup: Citizen Science – Existing efforts

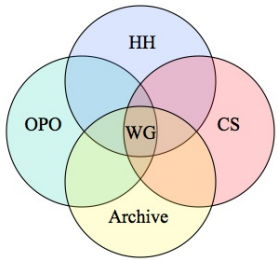
Zooniverse

Carol Christian and Alberto Conti worked closely with the Zooniverse people and COSMOS science team to make [Galaxy Zoo](#) happen.

In many cases this is the easiest way to get a project going.

In other cases it might be useful for STScI to get into the game ourselves to provide more flexibility (hookup a [basic version of “imexamine”](#) from IRAF to measure profile of objects – e.g., is this object a star or a star cluster?).

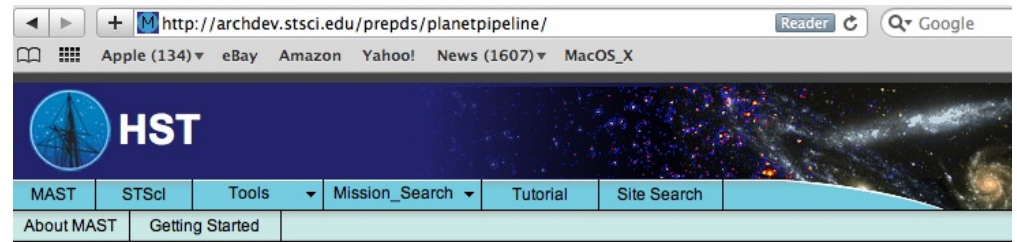
The screenshot shows the Galaxy Zoo: Hubble website. At the top, the browser address bar shows 'http://www.galaxyzoo.org/'. The website header features the 'GALAXY ZOO' logo with a glowing orange galaxy icon, and the word 'HUBBLE' in blue. A navigation menu includes links for Home, The Story So Far, How To Take Part, Classify Galaxies, Explore Galaxies, The Science, FAQ, Forum, Blog, and Contact Us. Below the navigation is a large image of a galaxy field. The main content area has a heading: 'Welcome to Galaxy Zoo, where you can help astronomers explore the Universe'. Below this is a paragraph explaining the project's goal and a 'Classifier Log In' section with a 'Click here to log in' button and links for 'Register' and 'Forgotten Password?'. There is also an 'Explore galaxies' section with a search bar and a 'Latest News' section.



Backup: Citizen Science – Existing efforts

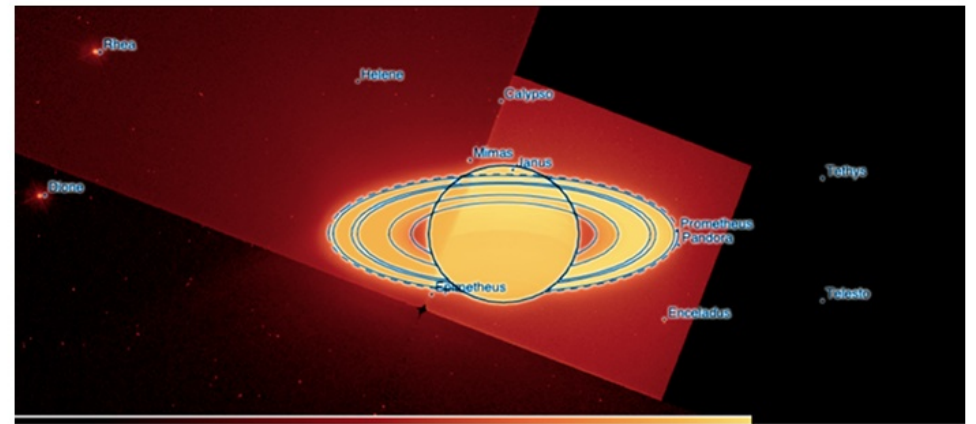
Max Mutchler’s archival program (ID 12142) to search their new “Planet Pipeline” images for planetary moons and other debris around planets.

They have partnered with Cosmic Quest to provide the infrastructure for the citizen science component of their project.



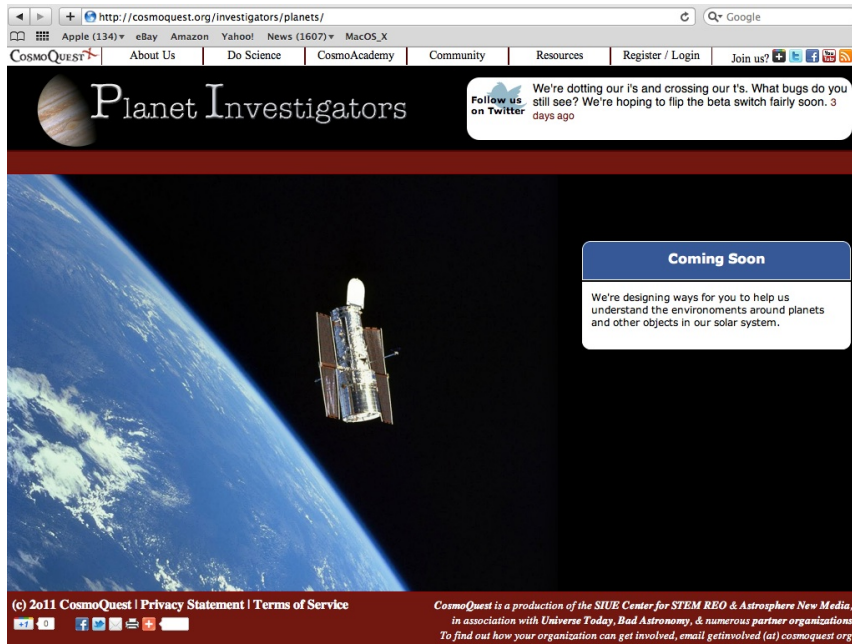
The Planet Pipeline: enabling data mining and citizen science with Hubble images of the Solar System

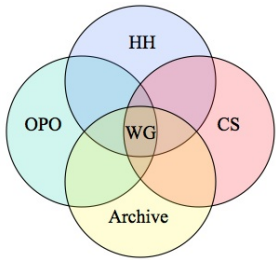
Max Mutchler, Michael H. Wong, Justin Higgins, Alberto Conti, Susana Deustua, Pamela L. Gay, Daniel Golombek, John Grunsfeld, Sean McKenna



In 15 years of service, the Wide Field and of Survey Instrument 2 (WFPC2) onboard the Hubble Space Telescope (HST) obtained over 10,000 frames of Solar System data. Since standard data reduction pipelines are typically not optimized for moving-target data, our “planet pipeline” will uniformly reprocess and catalog this WFPC2 image collection to make it more immediately science-ready. Some of our processing steps will utilize citizen scientists to perform visual inspections. Our corresponding database will enable robust queries which are more specific to planetary science, helping archival researchers quickly find and utilize the prepared images within our collection for a wide range of scientific analyses. We welcome suggestions (especially from veteran WFPC2 users) on the optimal treatment and organization of this data collection, and also to identify a broad range of analyses that might only be possible with visual inspections by citizen scientists.

Support for HST/AR program 12142 was provided by NASA through a grant from STScI, which is operated by the Association of Universities for Research in Astronomy, Inc., under NASA contract NAS 5-26555.





Backup: Citizen Science – Potential New Effort

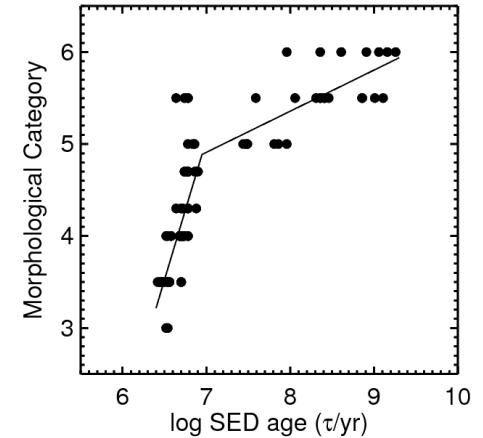
- Brad Whitmore is interested in setting up a Citizen Science project to help him measure the **ages of star clusters** using techniques developed in a recent paper published in the *Astronomical Journal*.
- He would use this resource for a number of datasets including the other 6 fields in M83 (ID = 11359), and future Hubble Heritage images.

THE ASTROPHYSICAL JOURNAL, 729:78 (14pp), 2011 March 10
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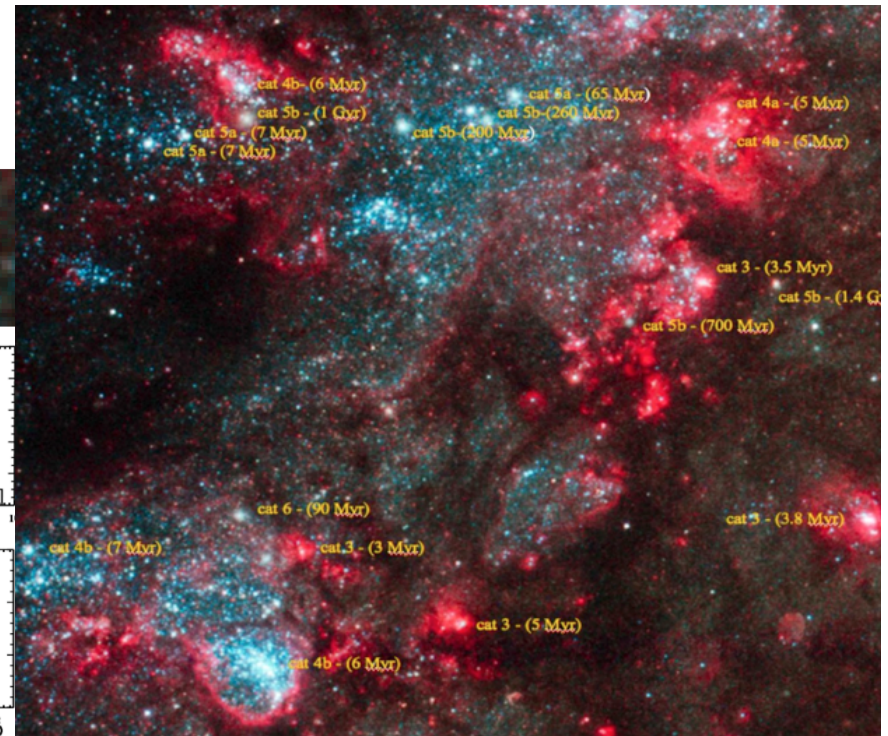
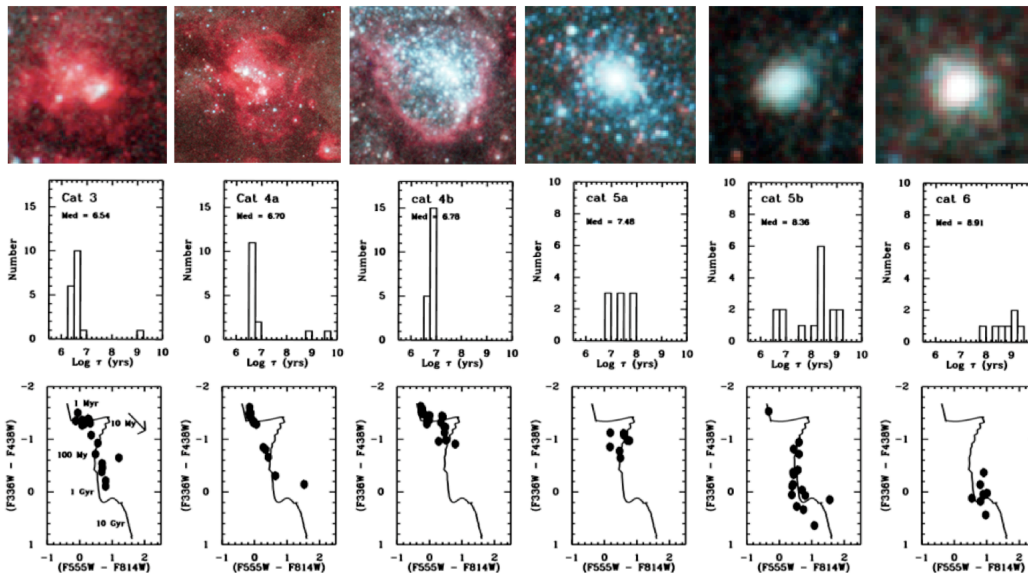
doi:10.1088/0004-637X/729/2/78

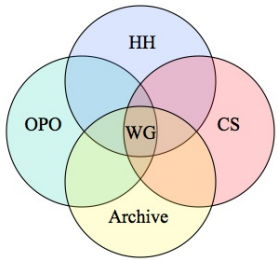
USING H α MORPHOLOGY AND SURFACE BRIGHTNESS FLUCTUATIONS TO AGE-DATE STAR CLUSTERS IN M83

BRADLEY C. WHITMORE¹, RUPALI CHANDAR², HWIHYUN KIM³, CATHERINE KALEIDA³, MAX MUTCHLER¹, MATT STANKIEWICZ¹, DANIELA CALZETTI⁴, ABHJIT SAHA⁵, ROBERT O'CONNELL⁶, BRUCE BALICK⁷, HOWARD E. BOND⁸, MARCELLA CAROLLO⁹, MICHAEL J. DISNEY⁴, MICHAEL A. DOPITA¹⁰, JAY A. FROGEL¹¹, DONALD N. B. HALL¹², JON A. HOLTZMAN¹³, RANDY A. KIMBLE¹⁴, PATRICK J. MCCARTHY¹⁵, FRANCESCO PARESCE¹⁶, JOSEPH I. SILK¹⁷, JOHN T. TRAUGER¹⁸, ALISTAIR R. WALKER¹⁹, ROGER A. WINDHORST⁴, AND ERICK T. YOUNG²⁰



category 3 emerging category 4a very young category 4b young category 5a interm/young category 5b intermediate category 6 old





Backup: Archive Lab Exercises

- HLA related materials are often **requested by educators at demos at AAS meetings.**
- HLA is starting an Educational Resource area: http://hlatest.stsci.edu/hla_edu.html.
- Working Group will facilitate archive/OPO collaborative projects.



Hubble Legacy Archive Educational Materials

. Color Magnitude Diagram

In this exercise students use real Hubble data to create a color magnitude diagram of M80, and then use theoretical isochrones to estimate its age and distance. The student should have some familiarity with stellar magnitudes and Hertzsprung-Russell diagrams, though details are given in the assignment. The full exercise requires knowledge of some form of coding (preferably IDL, to take advantage of the IDL Astronomy Library), but the simplified exercise only requires Microsoft Excel (or something similar). [Simplified Exercise](#)
[Original Exercise](#)

. V838 Monocerotis

In this exercise students use images of V838 Monocerotis taken several months apart to calculate the light echo's apparent rate of expansion. Students require access to software that allows them to view fits files (ex. DS9) [V838 Monocerotis Exercise](#)

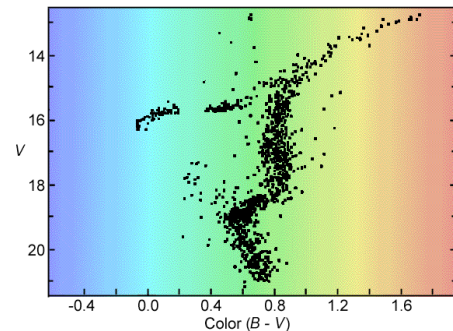
. M87 vs the Antennae

In this exercise students use color-color diagrams and single stellar population models (SSPs) to compare and contrast M87 and the Antennae galaxies. The full exercise is designed to be done using IDL, but it can be altered for lower-level students to only require excel. [M87 Exercise](#)

. Antennae Galaxies

In this exercise students use Hubble data to construct a color-color diagram of the clusters in the Antennae galaxies. This diagram is used in conjunction with a single stellar population model (SSP) and reddening vectors to identify a method of qualitatively age-dating some of the clusters. This exercise requires an understanding of astronomical magnitude systems (particularly Vegamag vs AB mag) and a basic understanding of extinction. In addition, it requires a good deal of coding (preferably IDL, to take advantage of the IDL Astronomy Library). [Age Dating Exercise](#)

Color Magnitude Diagram of M80



Background:

A familiar staple of the study of stellar evolution is the Hertzsprung-Russell (H-R) Diagram, like the one above. Developed in the early 1900s, the original H-R diagram plotted stars' spectral type versus their absolute magnitude. Today, diagrams plot stars' B-V color index (observed magnitude in the B band minus magnitude in the V band) versus magnitude. For obvious reasons, this is also known as a color magnitude diagram. H-R diagrams can also be made by plotting effective temperature versus luminosity. These two forms of the H-R diagram are similar in that they can both be used to identify stars of different ages and match observations to theoretical stellar evolution models, however the transformation between the two is not trivial. Because it relies solely on observation, we will focus on the color magnitude diagram (CMD).

A CMD of a single cluster gives a snapshot of that cluster's evolution, with the key features distinguishing groups of different ages. A star's mass governs both its initial position on the CMD and its evolution. Most stars fall along the very distinct main sequence, where they spend 90% of their lives. This stage is characterized by the fusing of hydrogen into helium in the stellar core. More massive stars burn fuel faster, and so exhaust their inner supply of hydrogen before their less-massive neighbors. When this occurs, these stars "turn off" the main sequence, brighten and migrate redwards.

Red giants are characterized by helium burning in the core, surrounded by a hydrogen shell. This helium burns into carbon, which combines with helium to make